# Perception: Motion Field Equations



#### Last Time:

#### Computing optical flow

$$\begin{pmatrix} \delta x^* \\ \delta y^* \end{pmatrix} = \left( \sum_{(x,y)} \nabla I_{t+1}(x,y) \nabla I_{t+1}(x,y)^T \right)^{-1} \left( \sum_{(x,y)} \Delta I_t(x,y) \nabla I_{t+1}(x,y) \right)^{-1} \left( \sum_{(x,y)} \Delta I_t(x,y) \nabla I_t(x,y) \nabla I_t(x,y) \right)^{-1} \left( \sum_{(x,y)} \Delta I_t(x,y) \nabla I_t(x,y) \nabla I_t(x,y) \right)^{-1} \left( \sum_{(x,y)} \Delta I_t(x,y) \nabla I_t(x,y) \nabla I_t(x,y) \right)^{-1} \left( \sum_{(x,y)} \Delta I_t(x,y) \nabla I_t(x,y) \nabla I_t(x,y) \nabla I_t(x,y) \right)^{-1} \left( \sum_{(x,y)} \Delta I_t(x,y) \nabla I_t(x$$





#### Hough Transform and RANSAC:





#### Repeat for $\boldsymbol{k}$ iterations

- 1. Choose a minimal sample set
- 2. Count the inliers for this set
- 3. Keep maximum, if it exceeds a desired number of inliers stop.

## Now: Computing 3D Velocity from 2D





## Now: Computing 3D Velocity from 2D





#### Which direction is the camera going?







Points from the original view







Moving along z-axis (optical axis)







Moving along z-axis (optical axis) Opposite way







Moving along x-axis







Moving along y-axis







Rotating along z-axis (optical axis)







Rotating along x-axis







Rotating along y-axis











Rotating along x-axis and y-axis, and translating





#### **Deriving Motion Field Equations**

To be able to predict and extract information from these optical flow fields, we need to derive a model of the equations of motion for points on an image











## Motion Model



### Motion Model



## Motion Model



## Motion Field Equations

The punchline is we can derive the motion field equations from these:

$$\dot{\mathbf{p}} = \frac{1}{Z} A(\mathbf{p}) \mathbf{V} + B(\mathbf{p}) \mathbf{\Omega}$$
$$\begin{pmatrix} u \\ v \end{pmatrix} = \frac{1}{Z} \begin{pmatrix} -1 & 0 & x \\ 0 & -1 & y \end{pmatrix} \begin{pmatrix} V_x \\ V_y \\ V_z \end{pmatrix} + \begin{pmatrix} xy & -(1+x^2) & y \\ 1+y^2 & -xy & -x \end{pmatrix} \begin{pmatrix} \Omega_x \\ \Omega_y \\ \Omega_z \end{pmatrix}$$

## Motion Field Equations - Derivation

Now to find the velocity on the image i.e. the motion field, we take the derivative of the 2D point:





## Motion Field Equations - Derivation

Substitute and rearrange:

$$\frac{\dot{\mathbf{P}}}{Z} - \frac{e_3^T \dot{\mathbf{P}}}{Z} \mathbf{p} = \frac{1}{Z} (I - \mathbf{p} e_3^T) \dot{\mathbf{P}}$$

From physics, velocity of a point in a rotating reference frame

$$\dot{\mathbf{P}} = -\mathbf{V} - \mathbf{\Omega} \times \mathbf{P}$$





## Solving the Motion Field Equations

We now know the form of the Motion Field Equations:

$$\dot{\mathbf{p}} = \frac{1}{Z}A(\mathbf{p})\mathbf{V} + B(\mathbf{p})\mathbf{\Omega}^{\prime}$$

How do we solve it

- Case 1: Known depth easy, simply use least squares
- Case 2: No translational velocity also least squares
- Case 3: No angular velocity slightly harder, but a form of least squares
- Case 4: Both translational and angular velocity with unknown depth fairly difficult



#### Case 1: Known Depth

With known depth the equations reduce to:

$$\dot{\mathbf{p}} = \frac{1}{Z} A(\mathbf{p}) \mathbf{V} + B(\mathbf{p}) \mathbf{\Omega} = \left(\frac{1}{Z} A(\mathbf{p}) \middle| B(\mathbf{p}) \right) \begin{pmatrix} \mathbf{V} \\ \mathbf{\Omega} \end{pmatrix}$$

And this can be solved with a least squares problem:

$$\mathbf{V}^*, \mathbf{\Omega}^* = \arg\min_{\mathbf{V},\mathbf{\Omega}} \sum_{i=1}^n \left\| \begin{pmatrix} \frac{1}{Z_i} A(\mathbf{p}_i) & B(\mathbf{p}_i) \end{pmatrix} \begin{pmatrix} \mathbf{V} \\ \mathbf{\Omega} \end{pmatrix} - \dot{\mathbf{p}}_i \right\|^2$$

THIS IS WHAT YOU WILL USE IN THE PROJECT You will need to run RANSAC to make this robust. How many points minimum do you need for this?



Vector of the equations reduce to a very simple form:  

$$\dot{\mathbf{p}} = B(\mathbf{p})\mathbf{\Omega}$$
And this can also be solved with a least squares problem:  

$$\Omega^* = \arg\min_{\Omega} \sum_{i=1}^n \|B(\mathbf{p}_i)\mathbf{\Omega} - \dot{\mathbf{p}}_i\|^2$$

## Case 3: No Angular Velocity

With no rotation the equations reduce to a different form than what we are used to:

$$\dot{\mathbf{p}} = \frac{1}{Z_i} A(\mathbf{p}) \mathbf{V} \iff [\dot{\mathbf{p}}]_{\times} A(\mathbf{p}) \mathbf{V} \stackrel{\scriptscriptstyle v_y}{=} 0$$

n

This can be reformulated as a minimum eigenvalue problem:

$$\arg\min_{\mathbf{V}:\|\mathbf{V}\|=1} \sum_{i=1}^{n} \|[\dot{\mathbf{p}}_{i}]_{\times} A(\mathbf{p}_{i})\mathbf{V}\|^{2}$$
$$= \sum_{i=1}^{n} ([\dot{\mathbf{p}}_{i}]_{\times} A(\mathbf{p}_{i})\mathbf{V})^{T} ([\dot{\mathbf{p}}_{i}]_{\times} A(\mathbf{p}_{i})\mathbf{V})$$
$$= \sum_{i=1}^{n} \mathbf{V}^{T} A(\mathbf{p}_{i})^{T} [\dot{\mathbf{p}}_{i}]_{\times}^{T} [\dot{\mathbf{p}}_{i}]_{\times} A(\mathbf{p}_{i})\mathbf{V}$$
$$= \mathbf{V}^{T} \left(\sum_{i=1}^{n} A(\mathbf{p}_{i})^{T} [\dot{\mathbf{p}}_{i}]_{\times}^{T} [\dot{\mathbf{p}}_{i}]_{\times} A(\mathbf{p}_{i})\right) \mathbf{V}$$



## Case 4: Everything Unknown

With no rotation the equations reduce to a different form than what we are used to:

$$\dot{\mathbf{p}} = \frac{1}{Z} A(\mathbf{p}) \mathbf{V} + B(\mathbf{p}) \mathbf{\Omega}$$

There is not too much we can do here as this is bilinear in our unknowns (depth and rotation). There are a few similar methods people have used to solve this:

- Method 1: Alternating minimization between V and  $\Omega/Z$
- Method 2: Marginalize out  $\Omega/Z$  and exhaustively search for



# Case 4: Everything Unknown

When we don't have any prior knowledge we are just left with the original equation:

$$\dot{\mathbf{p}} = \frac{1}{Z} A(\mathbf{p}) \mathbf{V} + B(\mathbf{p}) \mathbf{\Omega}$$

There is not too much we can do here as this is bilinear in our unknowns (depth and velocity). Velocity can only be recovered up to a scale. There are a few similar methods people have used to solve this:

- Method 1: Alternating minimization between V and  $\Omega/Z$  [1]
- Method 2: Marginalize out Ω/Z and exhaustively search for V[2]

[1] Optimal structure from motion: Local ambiguities and global estimates, Soatto and Brockett 1998

[2] Subspace methods for recovering rigid motion I: Algorithm and implementation, Heeger and Jepson 1992



$$\begin{array}{l} \textbf{Case 4: Everything Unknown} \\ \textbf{We can rewrite the equation as:} \\ \dot{\mathbf{p}}_{i} = \frac{1}{Z_{i}} A(\mathbf{p}_{i}) \mathbf{V} + B(\mathbf{p}_{i}) \Omega = \begin{pmatrix} A(\mathbf{p}_{i}) \mathbf{V} & B(\mathbf{p}_{i}) \end{pmatrix} \begin{pmatrix} \frac{1}{Z_{i}} \\ \Omega \end{pmatrix} \\ \Rightarrow \begin{pmatrix} \dot{\mathbf{p}}_{1} \\ \vdots \\ \dot{\mathbf{p}}_{n} \end{pmatrix} = \begin{pmatrix} A(\mathbf{p}_{1}) \mathbf{V} & \dots & 0 & B(\mathbf{p}_{1}) \\ \vdots & \ddots & \vdots & \vdots \\ 0 & \dots & A(\mathbf{p}_{n}) \mathbf{V} & B(\mathbf{p}_{n}) \end{pmatrix} \begin{pmatrix} \frac{1}{Z_{1}} \\ \vdots \\ \frac{1}{Z_{n}} \\ \Omega \end{pmatrix} \\ \Rightarrow \mathbf{u} = \begin{pmatrix} \mathbf{A}(\mathbf{V}) & \mathbf{B} \end{pmatrix} \begin{pmatrix} \rho \\ \Omega \end{pmatrix} \\ \Rightarrow \begin{pmatrix} \rho^{*} \\ \Omega^{*} \end{pmatrix} = \begin{pmatrix} \mathbf{A}(\mathbf{V}) & \mathbf{B} \end{pmatrix}^{+} \mathbf{u} \end{array}$$

## Case 4: Everything Unknown

With this we can solve for V through exhaustive search since V is in a bounded space (the sphere)

$$\operatorname{arg\,min}_{\mathbf{V}\in S^2} \left\| \mathbf{u} - \begin{pmatrix} \mathbf{A}(\mathbf{V}) & \mathbf{B} \end{pmatrix} \begin{pmatrix} \mathbf{A}(\mathbf{V}) & \mathbf{B} \end{pmatrix}^+ \mathbf{u} \right\|^2$$

A visualization of the loss function looks something like this:



0

#### Case 4: Everything Unknown A visualization of the loss function looks something like this:





#### Recap of Vision Thus Far



## **Basic Assumptions**

We take the thin lens model and approximate it as a pinhole model using calibration to make it good enough



$$x' = fx(1 + k_1r + k_2r^2 + ...) + x_0$$
  

$$y' = fy(1 + k_1r + k_2r^2 + ...) + y_0$$
  

$$\downarrow$$
  

$$x = \frac{X_c}{Z_c}, \quad y = \frac{Y_c}{Z_c}$$







## **Projective Transformations**



We learned about how to map points on the image to points on the image, specifically how planes transform

$$\lambda_{i} \begin{pmatrix} x'_{i} \\ y'_{i} \\ 1 \end{pmatrix} = \begin{pmatrix} h_{11} & h_{12} & h_{13} \\ h_{21} & h_{22} & h_{23} \\ h_{31} & h_{32} & h_{33} \end{pmatrix} \begin{pmatrix} x_{i} \\ y_{i} \\ 1 \end{pmatrix}$$
$$A = USV^{T} \implies h = V_{9}$$



## Pose from Projective Transformations



Assuming the points are on the plane we can solve for the exact pose from the transformation parameters and calibration

$$\begin{pmatrix} \hat{R}_1 & \hat{R}_2 & \hat{T} \\ \hat{R}_1 & \hat{R}_2 & \hat{T} \end{pmatrix} = \begin{pmatrix} \hat{r}_{11} & \hat{r}_{12} & \hat{t}_1 \\ \hat{r}_{21} & \hat{r}_{22} & \hat{t}_2 \\ \hat{r}_{31} & \hat{r}_{32} & \hat{t}_3 \end{pmatrix} = \begin{pmatrix} f & 0 & x_0 \\ 0 & f & y_0 \\ 0 & 0 & 1 \end{pmatrix}^{-1} \begin{pmatrix} h_{11} & h_{12} & h_{13} \\ h_{21} & h_{22} & h_{23} \\ h_{31} & h_{32} & h_{33} \end{pmatrix}$$
$$K^{-1} H$$

$$\begin{pmatrix} \hat{R}_1 & \hat{R}_2 & \hat{R}_1 \times \hat{R}_2 \end{pmatrix} = USV^T$$

$$R = U \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & \det(UV^T) \end{pmatrix} V^T$$

$$This arbit$$

$$T = \hat{T} \| \hat{R}_1 \|$$

This is arbitrary, can also use second column or average



#### PnP and P3P

If we have more general 2D-3D correspondences, we need to solve the problem more generally (V, V, Z)



Minimal case is P3P, which is helpful if we want to do RANSAC





### PnP and P3P

For P3P, we needed to solve a fairly complicated polynomial and go back to solve for the distances





#### Procrustes

Once we have the distances in P3P we can use that to solve for the position and orientation of the points

$$\arg\min_{R\in SO(3),T\in\mathbb{R}^{3}} \sum_{i} ||RP_{i} + T - P_{i}'||^{2}$$

$$\overline{P} = \frac{1}{N} \sum_{i} P_{i}, \ \overline{P'} = \frac{1}{N} \sum_{i} P_{i}$$

$$Z = \sum_{i} (P_{i} - \overline{P})(P_{i}' - \overline{P})^{T} = USV^{T}$$

$$R = V \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & \det(VU^{T}) \end{pmatrix} U^{T}$$

$$T = \overline{P'} - R\overline{P}$$
Can also use Procrustes

Can also use Procrustes in other contexts e.g. solving ICP





#### Procrustes

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## Epipolar Constraint

We also briefly went over how to get 3D pose from 2D-2D correspondences via the Essential matrix, and how to extract the pose from the essential matrix





## **Optical Flow**

Next we stepped back a bit and looked more at the image side and how to get correspondences, and some troubles that arise with that

$$(\delta x, \delta y) = \arg \min_{\delta x, \delta y} \sum_{(x,y) \in \mathcal{N}(x_0, y_0)} \left( \Delta I_t(x, y) - \nabla I_{t+1}(x, y)^T \begin{pmatrix} \delta x \\ \delta y \end{pmatrix} \right)^2$$

$$I_{t+1}(x + \delta x, y + \delta y)$$

$$\begin{pmatrix} \delta x^* \\ \delta y^* \end{pmatrix} = \left( \sum_{(x,y)} \nabla I_{t+1}(x, y) \nabla I_{t+1}(x, y)^T \right)^{-1} \left( \sum_{(x,y)} \Delta I_t(x, y) \nabla I_{t+1}(x, y) \right)$$

## **Outlier Rejection**

Given the high potential for error in these methods, we learned about methods to separate 'inliers' from 'outliers' using the Hough Transform, but primarily RANSAC  $\cos(\theta)x + \sin(\theta)y - \rho = 0$ 





#### Today: Motion Field Equations

And we learned how to take these things and derive motion models as opposed to merely static models, and we can estimate position and velocity

$$\dot{\mathbf{p}} = \frac{1}{Z} A(\mathbf{p}) \mathbf{V} + B(\mathbf{p}) \mathbf{\Omega}$$



#### Next Time

Now that we have estimates involving time. Given measurement error (even with something like RANSAC) there can still be estimation problems. Can we use temporal information to make things better? (Spoilers: yes). More importantly, how can we use temporal information in an optimal way while still being fast enough to be practical?



